# Towards per datum error characterization for radio occultation retrieval products

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#### Outline

- RO processing updates at JPL
- Approach in uncertainty estimates
- Results from CHAMP and COSMIC
- Summary

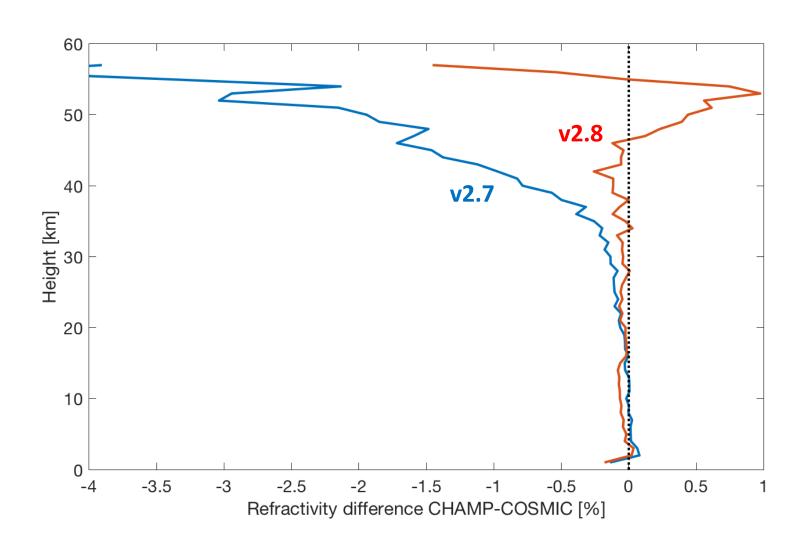
## RO processing at JPL

 Goal: a consistently processed RO data record from NASA/JPL receivers that includes CHAMP, SAC-C, GRACE, COSMIC, TSX, TDX, and KOMPSAT-5 from Level 0 to Level 3.

#### Recent changes:

- Cubic (previously quadratic) smoothing of phase to reduce biases above 20 km altitude where smoothing intervals are larger (v2.7).
- New Abel high-altitude initialization that aims to <u>reduce</u> <u>bias from noisy measurements</u> (which may impact consistency from different missions) and <u>reduce</u> <u>retrieval failures</u> (v2.8 in progress).

# CHAMP/COSMIC collocations < 300 km, 2 hr



## Motivation: Establishing GNSS RO as reference observations

- Following the GRUAN (GCOS Reference Upper Air Network) paradigm:
- ✓ Is traceable to an SI unit or an accepted standard
- ✓ Provides a comprehensive uncertainty analysis
- ✓ Is documented in accessible literature
- ✓ Is validated (e.g. by intercomparison or redundant observations)
- ✓ Includes complete meta data description
- ✓ Important to distinguish contributions from systematic error and random error

## Some existing works

#### Kursinski et al. 1997

 Comprehensive theoretical analysis with multiple error sources.

#### Kuo et al. 2005

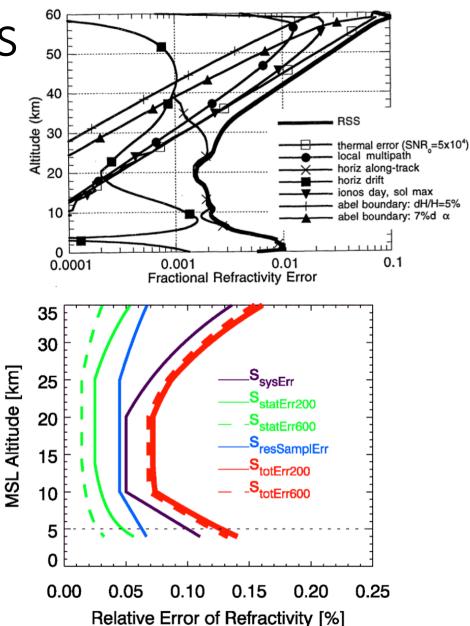
 Derived error estimates based on actual retrieval comparison with NWP forecasts.

#### Scherlin-Pirscher et al. 2011

 Explicit separation of systematic and random errors, plus sampling error for climatological averages.

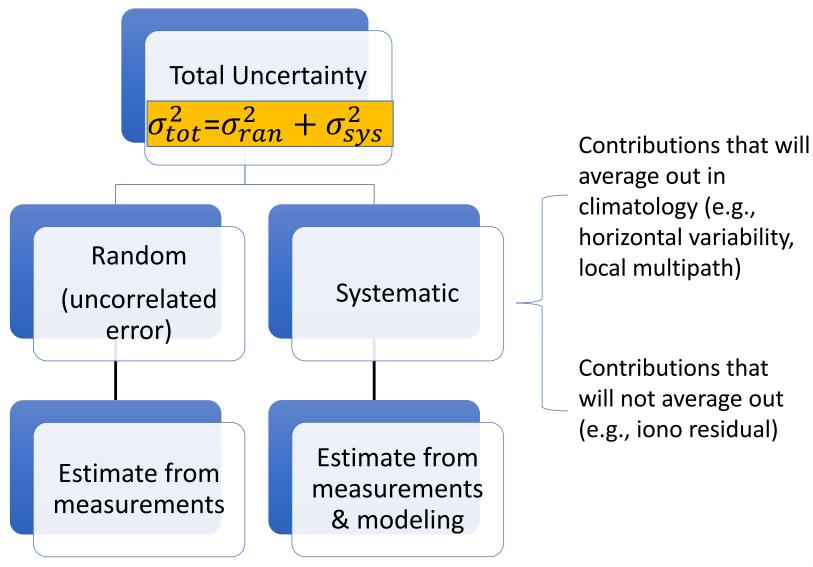
#### Schwarz et al. 2017

- Detailed error estimate and propagation.
- Similar objectives as ours.



Independent uncertainty estimates specific to a retrieval system are desirable.

# Uncertainty estimation (separate random & systematic)



## Random errors: Bending angle

Estimate phase noise from the L1 and L2 excess phase data:

- Detrend phase and compute standard deviation over 1 sec to get the 1-sec phase noise.
- 2. Scale to actual smoothing interval *T*-sec if needed.
- 3. Derive bending angle uncertainty using the following expressions [Hajj et al. 2002]:

$$\sigma_i(M) \approx \frac{c}{f_i V} \left( \frac{\nu \sigma_\phi}{\Delta t \ M^{3/2}} \right)$$

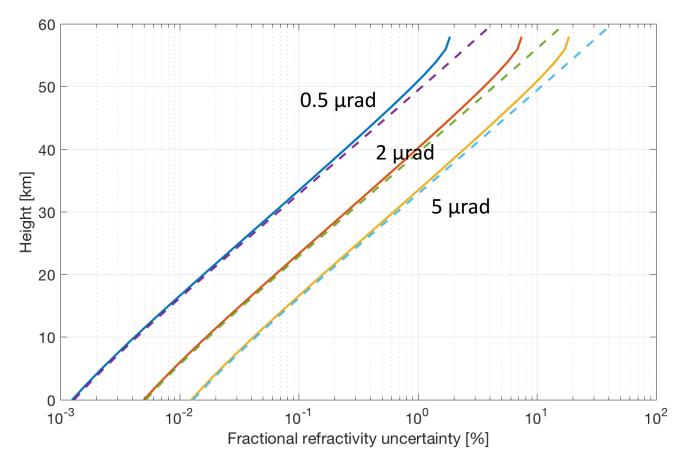
 $\Delta t$  = sample time (e.g., 20 msec) M = number of data points in the smoothing interval (e.g., 50) V = tangent point velocity (e.g. 2 km/s)

$$\sigma_n^2 = \sigma_1^2(M_1) + (1.54)^2 \left[ \sigma_1^2(M_2) + \sigma_2^2(M_2) \right]$$
 Coarse smoothing (M<sub>2</sub> > M<sub>1</sub>)

## Random errors: Refractivity

$$\sigma_j^{(N)} = \left[\sum_{i=j+1}^M F_{ji}^2 \sigma_i^2\right]^{1/2}$$

where 
$$F_{ji}=rac{10^6}{\pi}rac{\delta a_i}{\sqrt{a_i^2-a_j^2}}$$



**Solid lines**: BA contribution from impact height < 60 km

**Dashed lines**: impact height < 80 km

### Sources of systematic BA errors

Not an exhaustive list!

#### 1. Residual ionosphere

- 2. Horizontal inhomogeneity
- 3. Local multipath
- 4. POD (pos, vel, clock)

#### For lower troposphere:

- 5. Tracking error? [Zus et al. 2014]
- 6. Retrieval nonlinearity? [Sokolovsiy et al. 2010]

## Systematic errors: Refractivity

From <u>systematic error of BA</u>:

$$\langle \Delta N_j \rangle = \sum_{i=j+1}^{M} F_{ji} \langle \Delta \alpha_i \rangle$$

• Abel Upper Boundary (UB) condition introduces uncertainty in refractivity. For exponential extrapolation above  $a_u$ , we estimate the refractivity uncertainty at  $a_j$  below  $a_u$  due to scaleheight H uncertainty as

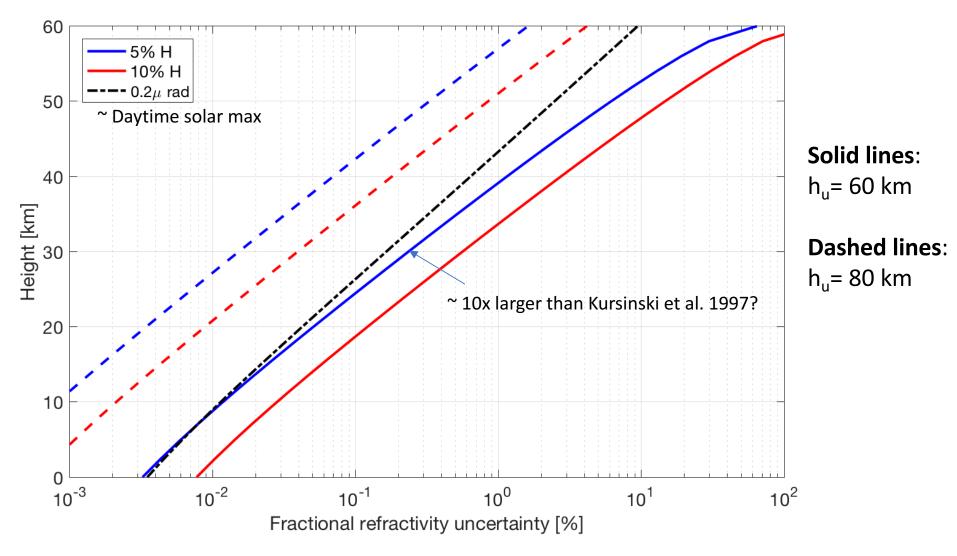
$$\langle \Delta N_j \rangle^{(U)} = U_j(H \pm \Delta H) - U_j(H)$$

where *U* is given by [Gleisner and Healy, 2013]

$$U(a; H) \approx 10^6 \alpha_u e^{-(a-a_u)/H} \sqrt{\frac{H}{2\pi a}} \operatorname{erfc}\left(\sqrt{\frac{a_u - a}{H}}\right)$$

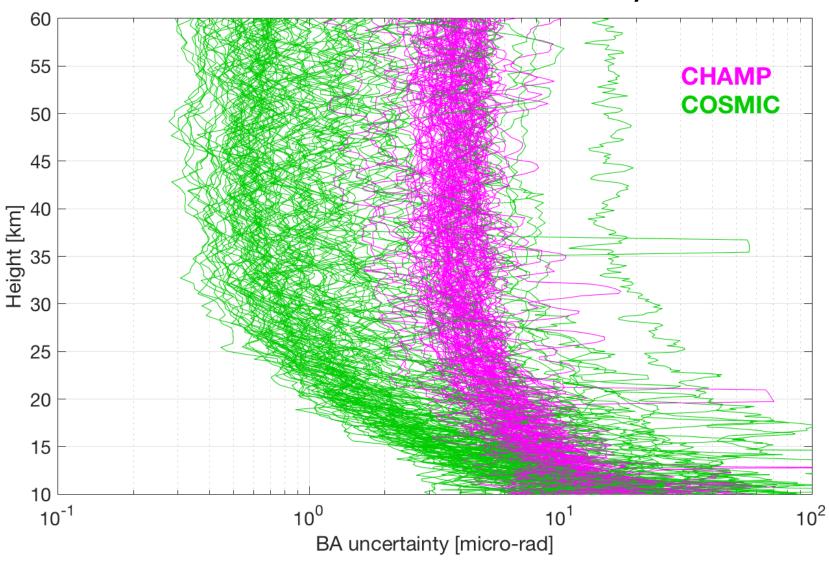
ΔH will be determined based on residuals to each fit

#### Iono & UB errors

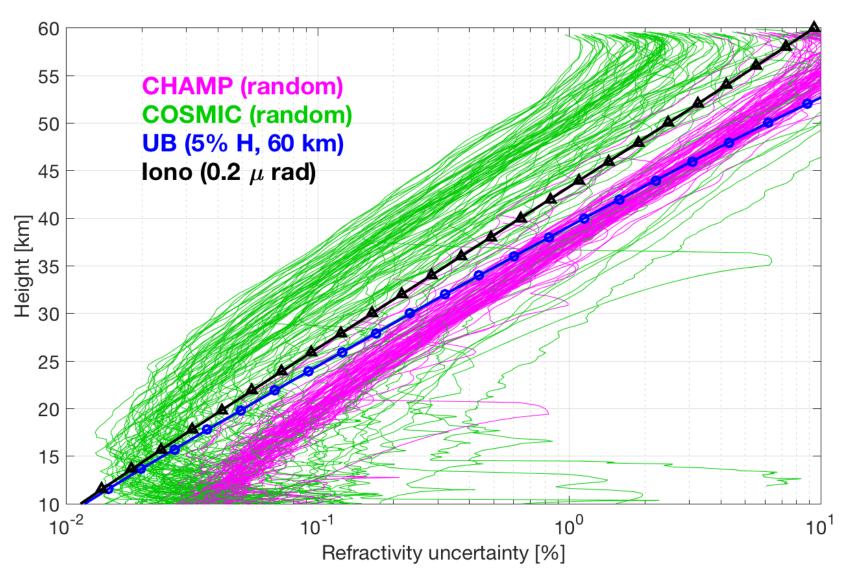


## Examples from CHAMP & COSMIC

#### **Estimated random BA uncertainty**

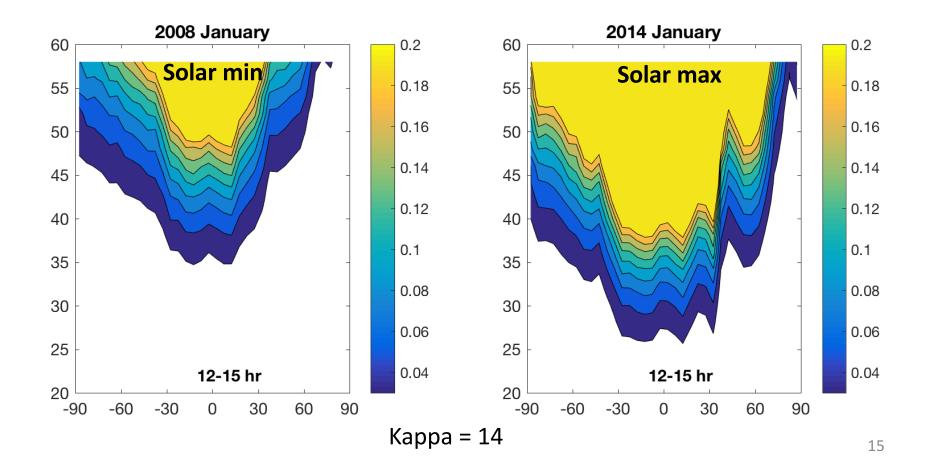


## Examples from CHAMP & COSMIC



#### Better estimate of iono error

$$\begin{split} \alpha_{\rm c}(a) &= \alpha_{\rm L1}(a) + \frac{f_2^2}{f_1^2 - f_2^2} (\alpha_{\rm L1}(a) - \alpha_{\rm L2}(a)) \\ &+ \kappa(a) (\alpha_{\rm L1}(a) - \alpha_{\rm L2}(a))^2, \end{split}$$
 Healy and Culverwell, 2015



### Summary

- Progress towards per datum uncertainty characterization of RO retrieval products at JPL.
- A few dominant error sources have been considered so far.
- Uncertainty estimates need to be verified (through comparisons with other data, RO pairs, etc.) and refined.
- Per datum uncertainty gives an effective approach in quality control.